

# **Euclid: an ESA mission to map the Dark Universe**

**Presentation to the Portuguese  
Delegation and Industry**

**Lisbon, March 19<sup>th</sup> 2012**

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# Outline of the presentation



## 1. EUCLID Science Objectives

2. EUCLID Mission and Payload Module Overview

3. Service Module Overview

4. Schedule

## Search for the critical density

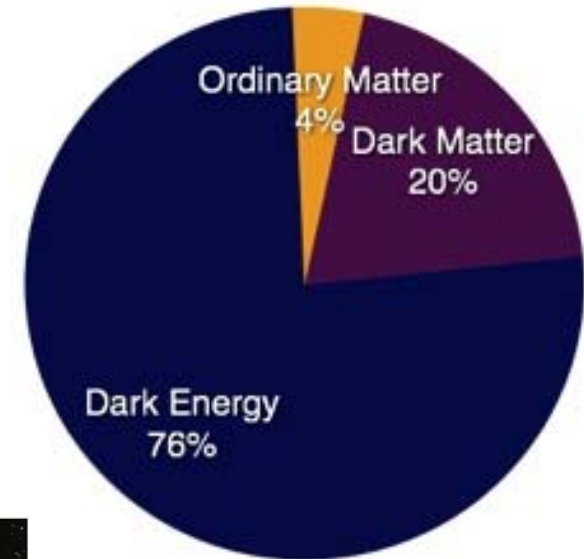
- Our Universe is observed to be **Euclidean or “flat”**, best described with the **Big-Bang inflation model**. This is confirmed by the CMB observation
- **The flatness requires a critical density ( $\Omega=1$ ):**
  - Amount of protons+neutrons (baryons) is constrained by direct observations and the theory of nucleosynthesis
    - **Baryons must be less than 3-5% of critical density**
  - Dark Matter is needed which is “Cold”, i.e. the dark matter particles have negligible velocity, but observations show
    - **Cold DM cannot be more than 10x baryonic mass**
  - Remaining missing mass not found, instead, the *cosmological constant*  $\Lambda$  needed re-introduction
    - **A positive  $\Lambda$  can explain  $\sim 80\%$  of the critical density**
- **Supernova observations indicate accelerated expansion of the Universe => Cosmological Constant**

# EUCLID Scientific Objectives (2/6)

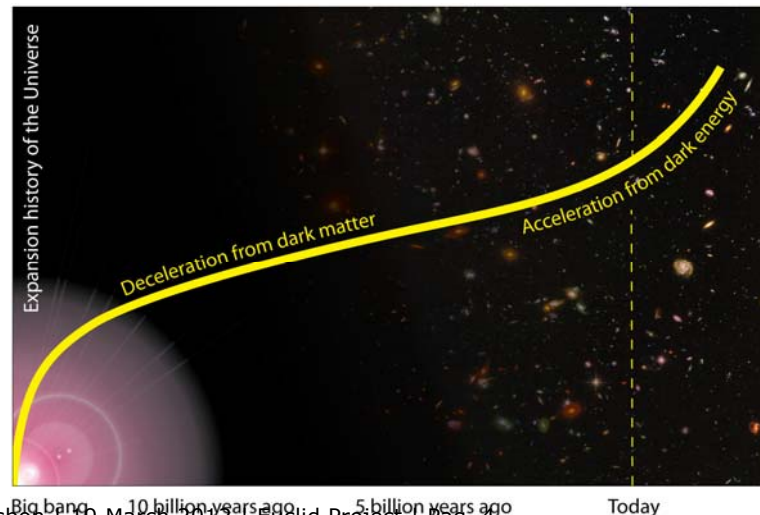
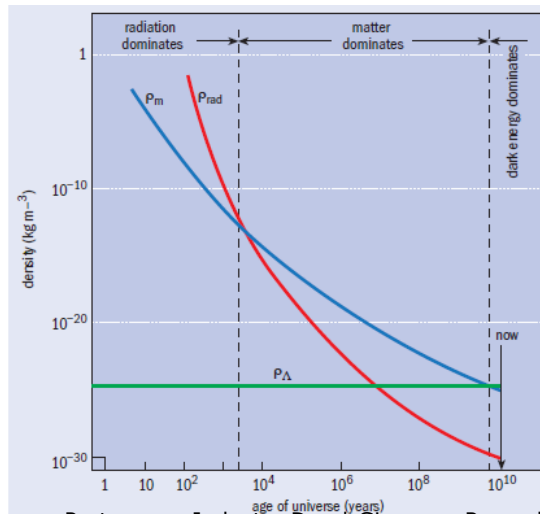


## The Concordance model: Lambda Cold Dark Matter - $\Lambda$ CDM

- $\Lambda$ CDM is so far the only model within the framework of General Relativity which fit all data at once, including
  - Nucleosynthesis of light elements
  - Cosmic Microwave Background measurements
  - Accelerating expansion of the Universe, from Supernovae observations
  - Cosmic structures, galaxy clustering
- $\Lambda$ CDM does not explain the physics, it leaves open several fundamental questions, in particular the properties and nature of the dark components



Concordance Model



# EUCLID Scientific Objectives (3/6)



## Summary

Issue	Euclid's Targets
<b>What is Dark Energy</b>	<b>Measure the Dark Energy equation of state parameters</b> $w_p$ and $w_a$ to a precision of 2% and 10%, respectively, using both expansion history and structure growth.
<b>Beyond Einstein's Gravity</b>	<b>Distinguish General Relativity from modified-gravity theories</b> , by measuring the galaxy clustering growth factor exponent $\gamma$ with a precision of 2%.
<b>The nature of dark matter</b>	<b>Test the Cold Dark Matter paradigm</b> for structure formation, and measure the sum of the neutrino masses to a precision better than 0.04eV when combined with Planck.
<b>The seeds of cosmic structure</b>	<b>Improve by a factor of 20 the determination of the initial condition parameters</b> compared to Planck alone. $n$ (spectral index), $\sigma_8$ (power spectrum amplitude), $f_{NL}$ (non-gaussianity)

## Concept

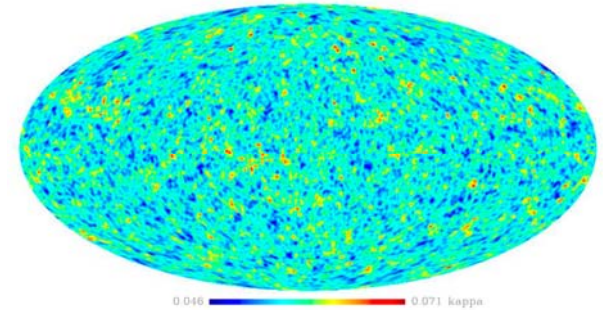
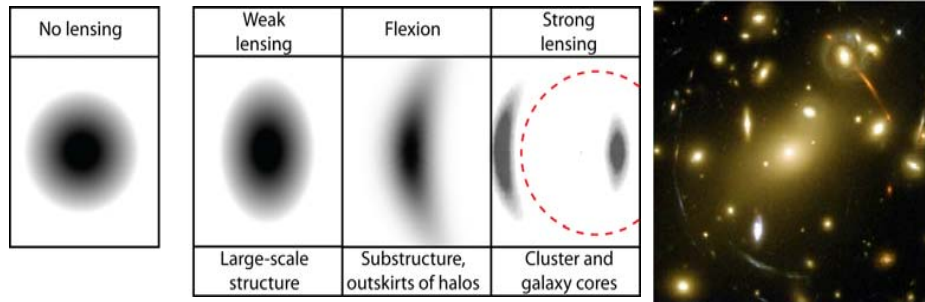
- Euclid will perform a sky survey and is optimised to measure simultaneously two principal dark energy probes:
  - **Weak Lensing**
  - **Baryon Acoustic Oscillations (galaxy distribution power spectrum)**
  
- **Science Requirements**
  - Minimum survey area of 15,000 deg<sup>2</sup> (~1/4 of sky)
  - Determine the shapes and shear of statistical samples of galaxies with a density of 30-40 galaxies/arcmin<sup>2</sup>. Total of 1.5 billion galaxies
  - Determine the photometric redshifts of the weak lensing galaxies with  $dz/(1+z)=0.05$  down to 0.03
  - Measure spectroscopic redshifts with  $dz/z < 0.001$  in the same volume.

**The mission is named in honour of the pioneer of geometry.**

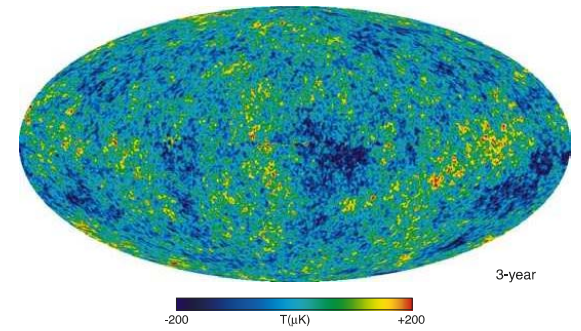
# EUCLID Scientific Objectives (5/6)



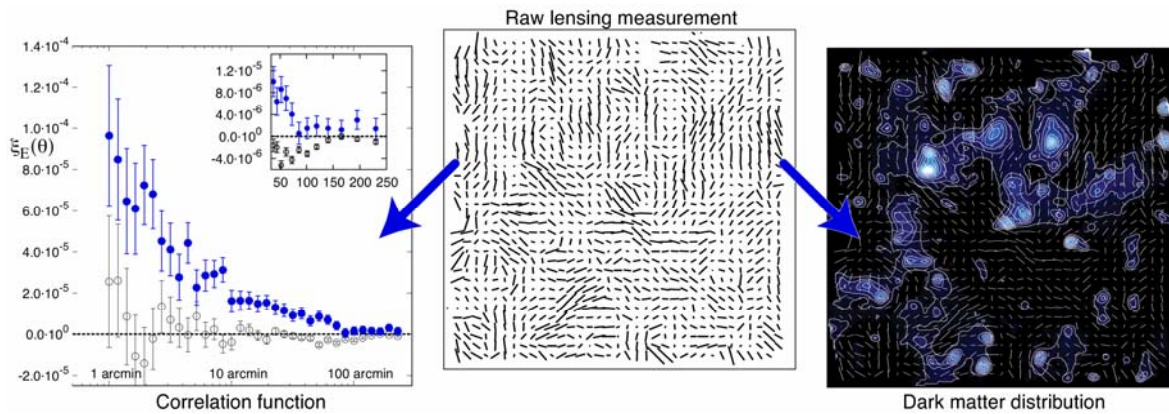
## Weak Lensing Probe



Weak Lensing (simulation) result:  
Dark Matter distribution at  $z=2$



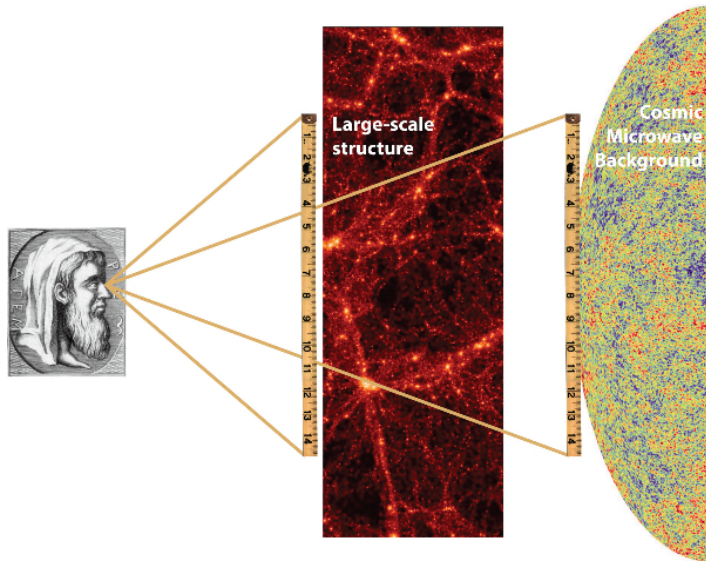
WMAP CMB  
Structure at  $z=1100$



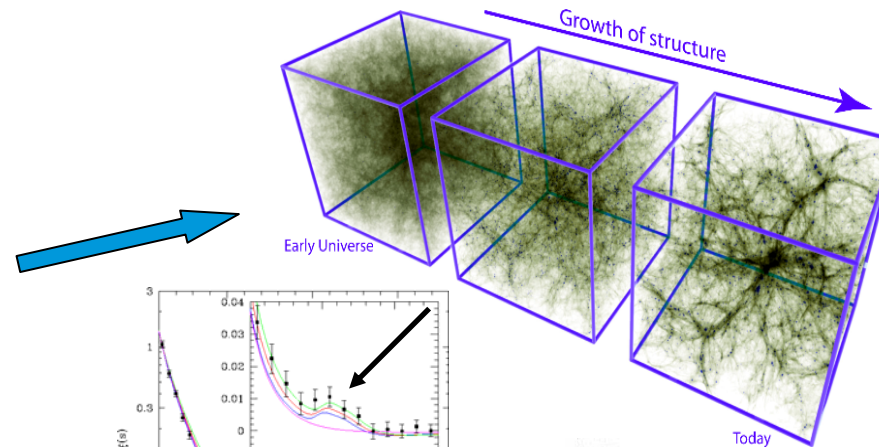
# EUCLID Scientific Objectives (6/6)



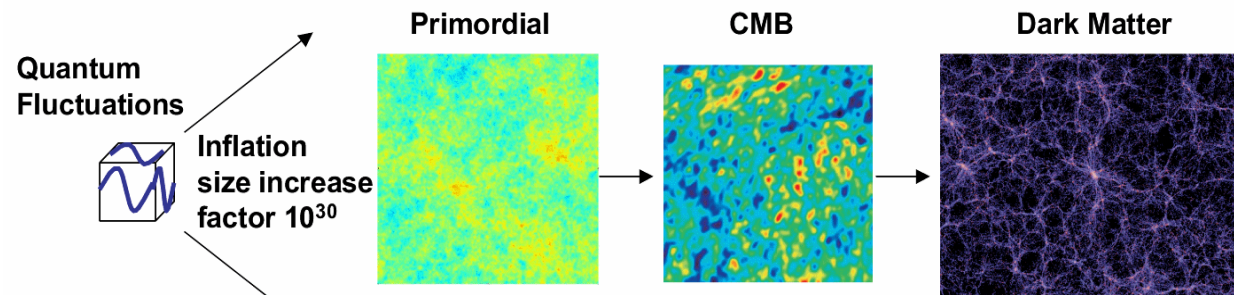
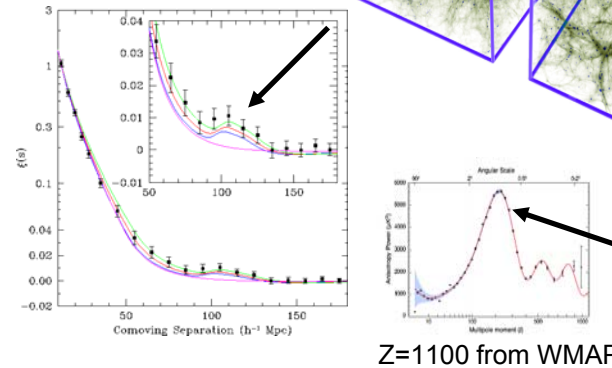
## Baryon Acoustic Oscillations



Baryon Acoustic Oscillations: measure typical scales of galaxy structures in comoving frame



BAO will map the cosmic web using baryon distribution



**Determination of Initial Conditions**

pace Agency



# Outline of the presentation



1. EUCLID Science Objectives

**2. EUCLID Mission and Payload Module Overview**

3. Service Module Overview

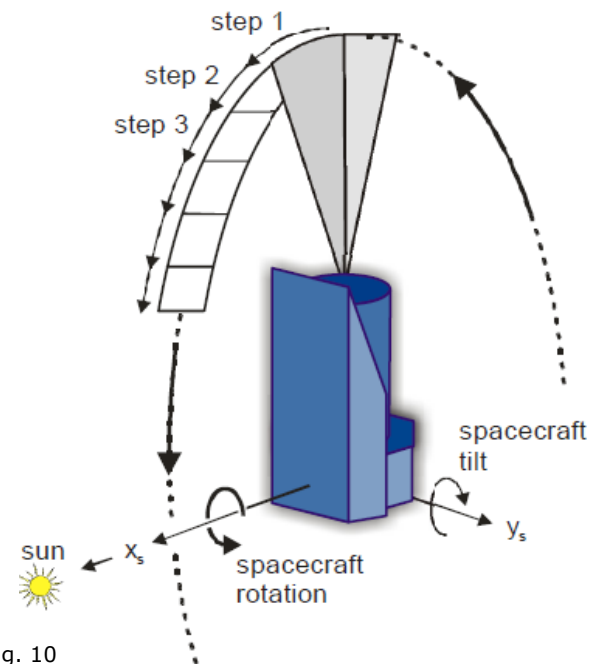
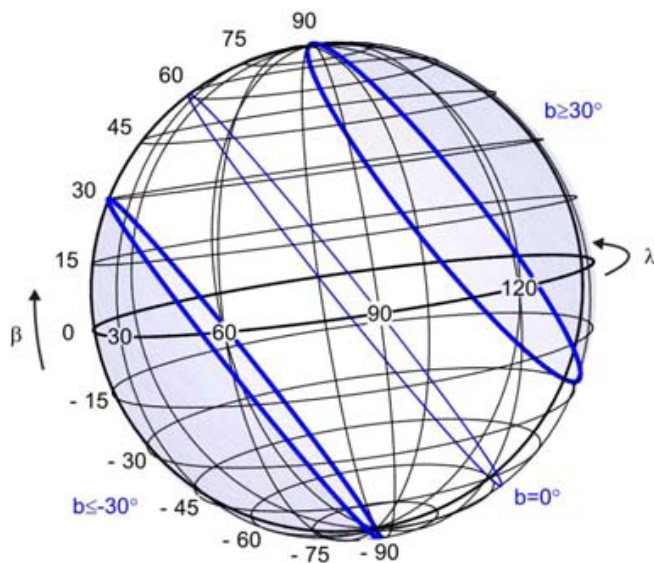
4. Schedule

# EUCLID mission overview (1/8)



## Mission facts

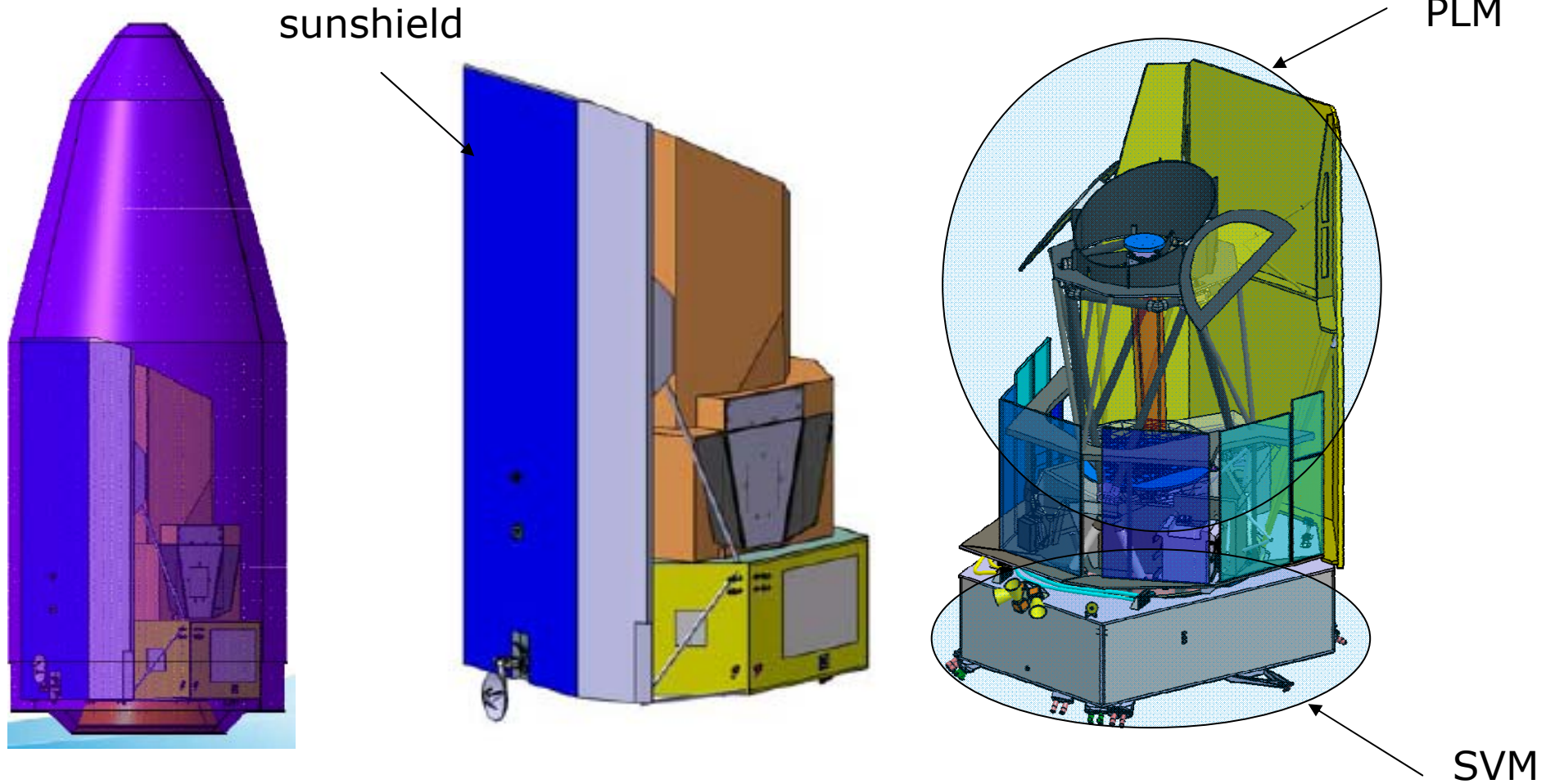
- Survey mission with 6.25 year nominal operation duration.
- The Spacecraft will be launched by a Soyouz ST2-1B from French Guyana to L2, maximum launch mass is 2160 kg.
- The 3 axis stabilized spacecraft is operated in step and stare mode (around the S/C sun axis) to observe the extra galactic latitudes  $> 30$  degrees.



# EUCLID mission overview (2/8)



## Quick view of the spacecraft



Astrium config (PM6)

TAS config (PM5)

## Payload Module facts

- The EUCLID payload module consists of:
- Optical configuration: 3-mirrors anastigmatic Korsch
  - field of view:  $0.763 \times 0.709 \text{ deg}^2$ , 0.45deg off-axis
  - free aperture: 1.2 m<sup>2</sup>
  - WFS and M2 mechanism for calibration
  - optical quality: WFE  $\leq 70\text{nm rms}$  (NISP channel)
  - common bench for telescope and instruments, interface to SVM
  - Zerodur/carbon or Silicon Carbide SiC-100 technologies
  - Dichroic at the exit pupil transmit IR light to NISP and reflects visible light to VIS
  - A visible imager (VIS)
  - A near-IR instrument (NISP)

## VIS Instrument

A visible imager:

- limiting magnitude:  $\text{mag}_{AB}=24.5$
- redshift resolution:  $z/(1+z)$  3%–5%
- spectral range : 550–920nm
- focal plane: 6 × 6 CCDs (e2v, 12 × 12 m<sup>2</sup> pixels, 4096 × 4096 pixels)
- plate scale: 0.1arcsec/pix
- field of view:  $\text{FoV}=0.787 \times 0.709\text{deg}^2$
- focal length:  $f=24.5$  m

## NISP Instrument

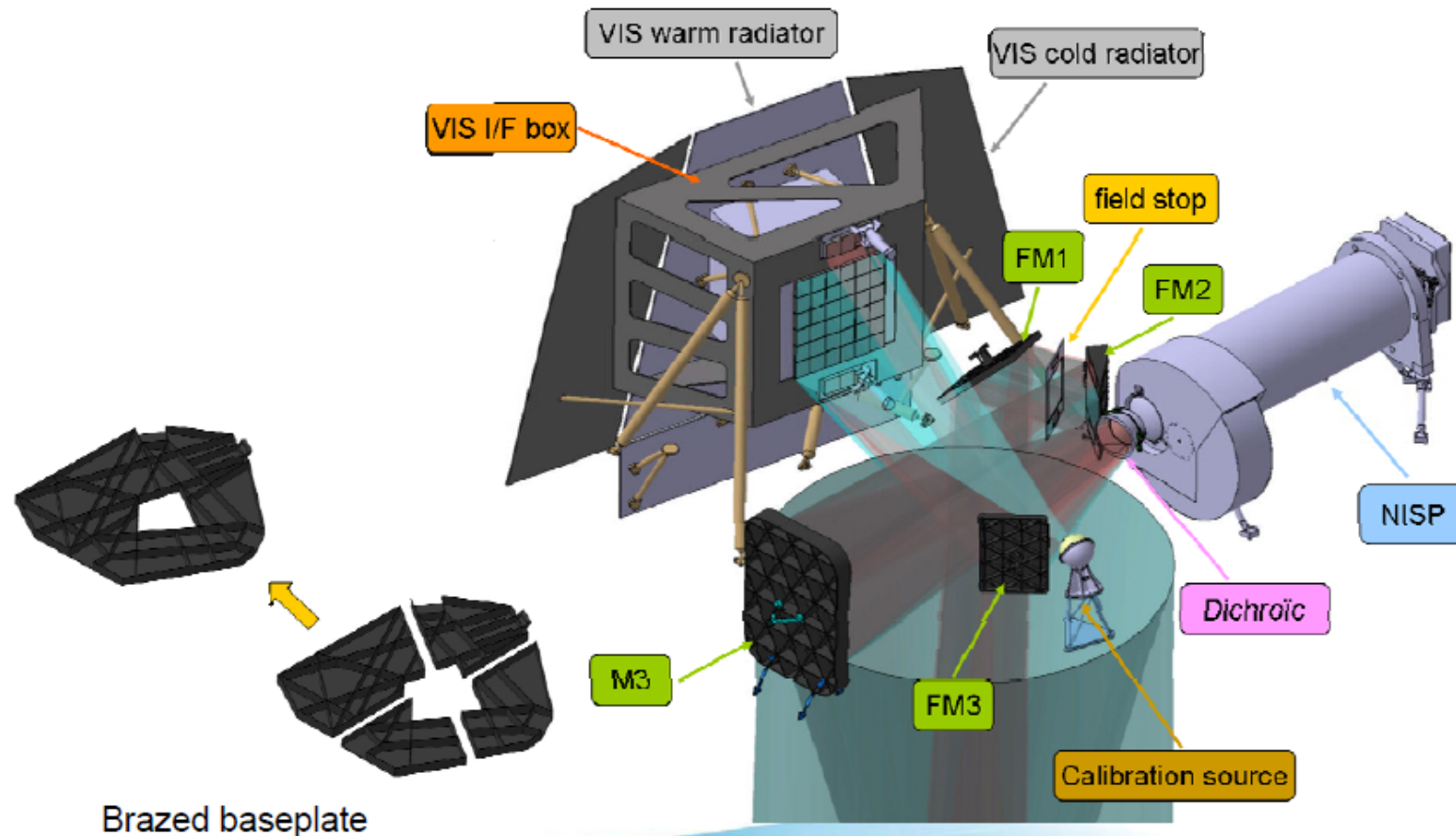
A near-infrared spectrograph photometer:

- Photometry
  - limiting magnitude: magAB=24.5
  - redshift resolution:  $z/(1+z)$  3%–5%
  - spectral range : 920–1146nm (Y), 1146–1372nm (J), 1372–2000nm (Hp)
- Spectroscopy
  - limiting magnitude: magAB=19.5
  - redshift resolution:  $z/(1+z)$  0.1%
  - spectral range : 1000–2000nm
  - spectroscopic resolution:  $\sim$  =500
- focal plane: 4 4 Hawaii 2RG (Teledyne, 2048 2048pixels, 2.4 m)
- plate scale: 0.3arcsec/pix
- field of view: FoV=0.763 0.722deg<sup>2</sup>
- focal length:  $f\sim 6.1$  m

# EUCLID mission overview (6/8)



## Payload facts

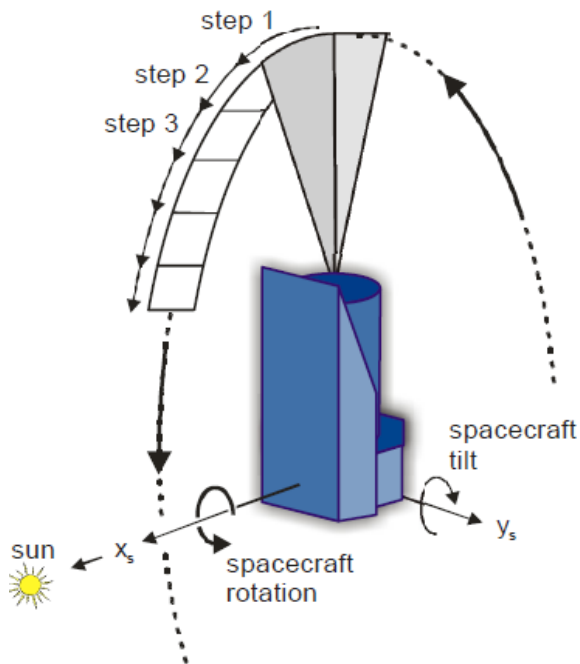


Braze baseplate  
Astrum configuration  
(passive thermal control, SiC techno)

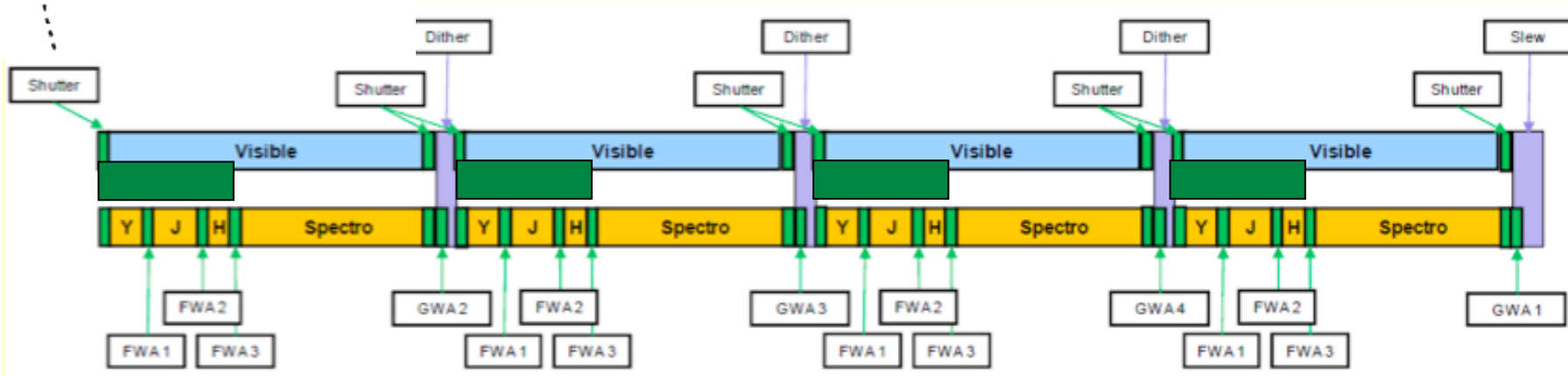
# EUCLID mission overview (7/8)



## Typical operation



- For each step (field), 3 dithers are performed at Spacecraft level leading to a total of 4 frames.
- For each dither the 3 photometric bands are observed thanks to the NI-FWA actuation
- For spectroscopy, a combination of 2 filters and 2 grisms (identical but different dispersion directions) is used for the 4 frames
- VIS is integrating in parallel with the spectro to avoid any tilt disturbances from NI-FWA and NI-GWA
- VIS shutter is kept closed during photometry





# EUCLID mission overview (8/8)



## Ground Segment

1. Mission Operation Centre
  - a. at ESOC (Darmstadt, Germany)
  
2. Science Operation Center
  - a. at ESAC (Villafranca, Spain)
  
3. Ground Stations:
  - a. Cebreros and Malargue antennas
  - b. Daily science communication:
    - ~ 850 Gbits in K band (26 GHz)
  - a. Command and control in X band



# Outline of the presentation



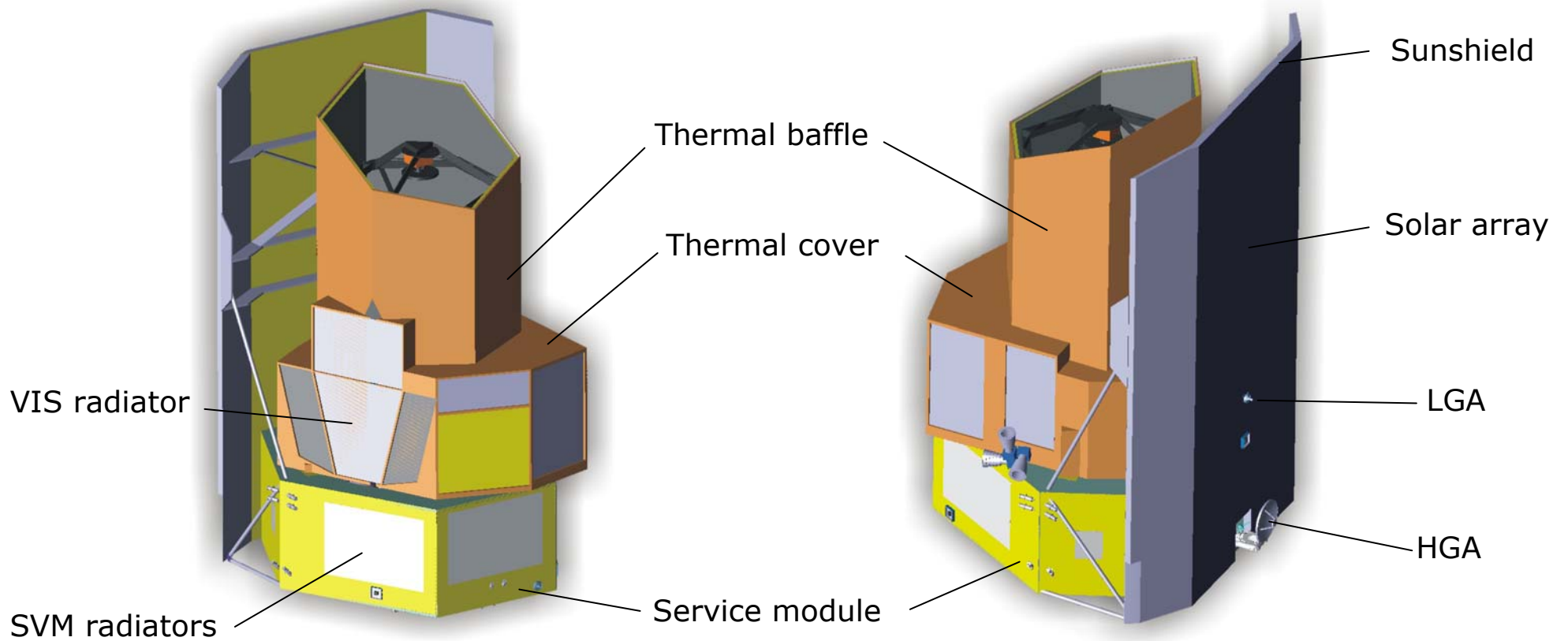
1. EUCLID Science Objectives

2. EUCLID Mission and Payload Overview

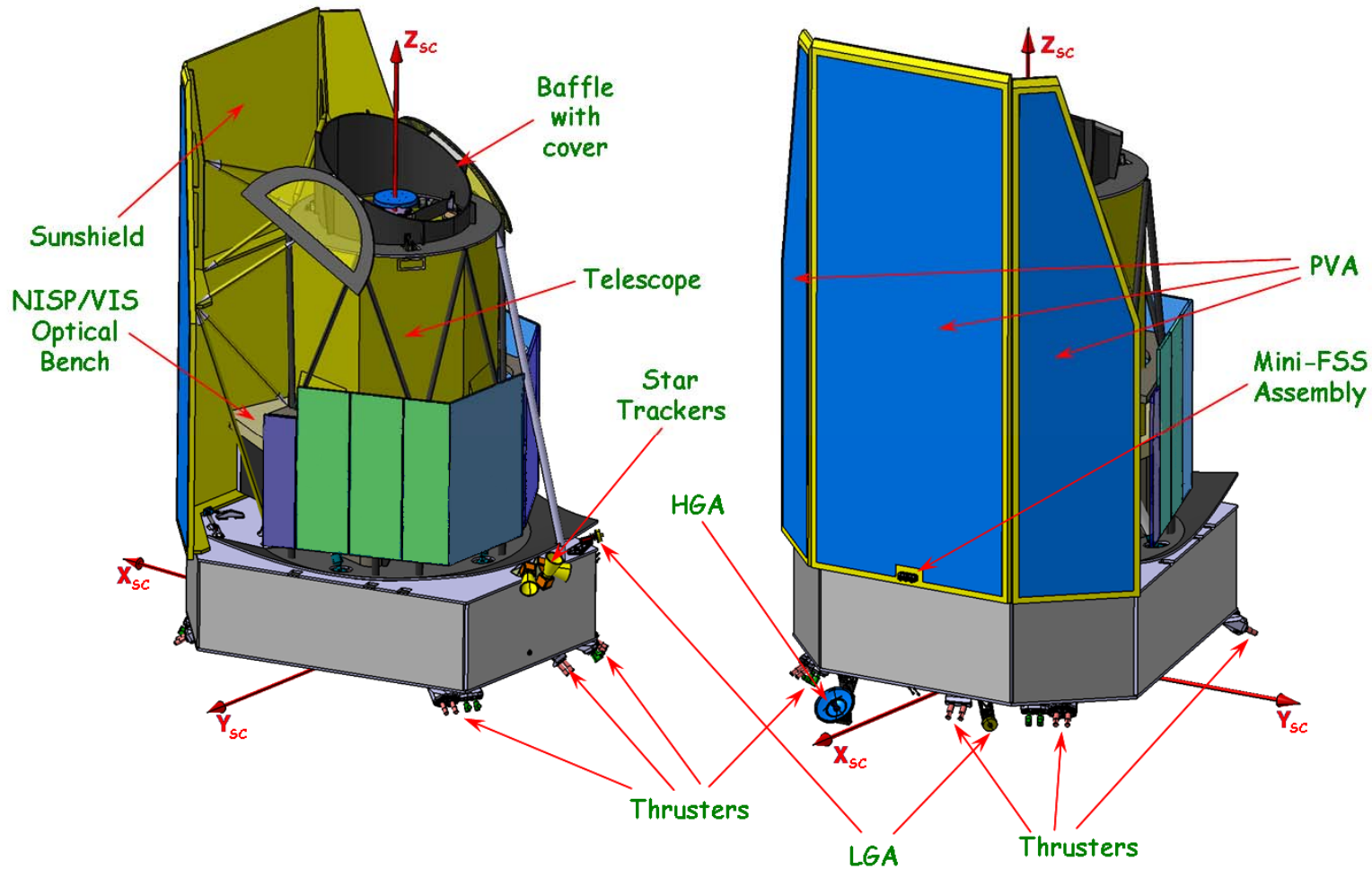
**3. Service Module Overview**

1. Schedule

## Astrium concept



## TAS concept



## ASTRIUM concept

### Telescope

1. Primary Mirror: SiC
2. Cold Telescope ( $T \sim 150\text{K}$ )
3. Passive Thermal Control

### AOCS

1. Fine pointing: Cold Gas + FGS & Gyro
2. Slews: Cold Gas + Star Tracker & Gyro

## THALES concept

### Telescope

- Primary Mirror: Zerodur
- Cold Telescope ( $T \sim 240\text{K}$ )
- Active Thermal Control

### AOCS

- Fine pointing: Cold Gas + FGS & Gyro
- Slews: Reaction Wheel + Star Tracker & Gyro

# Euclid Service Module Overview (1/4)



## Attitude and Orbit Control Subsystem

- Sensors:
  - fine guidance sensor (half frame transfer CCDs, 1Hz bandwidth, close to VIS focal plane)
  - inertia measurement unit
  - high performance star tracker
  - fine sun sensors
- Micro-propulsion
  - science mode manoeuvres (field step, dither step for AST), attitude control (solar pressure compensation)
  - proportional cold-gas, 6 thrusters (1000 N, cold redundant)
- Reaction Wheels for slew and dither steps (only TAS)
- Chemical propulsion
  - orbit maintenance, transfer correction, safe mode
  - mono-propellant hydrazine (blow-down mode), membrane tank, 6 thrusters (20N, cold redundant)
- Pointing performance
  - RPE < 15/15/150 mas rms per x/y/z-axis (700s)
  - dither step: < 75 s
  - field step: < 350 s

# Euclid Service Module Overview (2/4)



## Communication Subsystem

- Science data 850 Gbit/day
- Memory 4 Tbit Non volatile flash memory, 3 days storage capability
- Link window 3.5h/day effective downlink time, Cebreros ground station ( 35m)
- TM/TC
  - 2 X-band LGA,
  - 2 resp. 5 kbit/s downlink,
  - 4 resp. 128kbit/s uplink,
  - hot redundant receiver and cold-redundant transmitter
- Science telemetry
  - K-band (26GHz),
  - steerable HGA, 40cm,
  - 75Mbit/s downlink,
  - GMSK modulation,
  - internally cold-redundant

# Euclid Service Module Overview (3/4)



## Electrical and Power Subsystem

- Electrical architecture:
  - OBC (e.g. LEON)
  - SSMM 4+2 Tbit
  - packet telemetry/telecommand (PUS compatible)
  - MIL-STD-1553B bus for PLM and SVM unit interfaces
  - SpaceWire link to EIU
- Power subsystem:
  - 28V regulated,
  - MPPT,
  - 75Ah Li-Ion battery
- Solar array:
  - body mounted (on sunshield),
  - GaAs triple junction cells (28% efficiency)
  - 9.8m<sup>2</sup>



# Euclid Service Module Overview (4/4)



## Structure and Thermal Subsystems

- Service Module:
  - hexagonal shape, central cone, 6 shear walls, 6 hard points for PLM interface
  - Al honeycomb, CFRP facets
  - 1666mm SF Soyuz launcher interface
  - radiative cooling, foil radiators
- Unit accommodation
  - position of tanks optimised for science mode manoeuvres
  - high dissipation units on cold side
- Sunshield
  - stiff design (angled wings, ribs, struts)
  - Al honeycomb, CFRP facets
  - planar solar array mounted on Sun-facing side

# Euclid Industrial Schedule



## Astrium version

